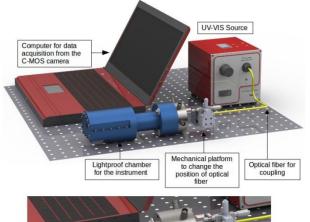
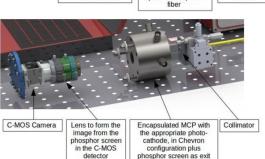
Financiang from the Mexican Space Agency



- Development of a computational model of Microchannel plates based on the physics involved.
- Generation of the visible photon shower in the output of the MCP
- Evaluation of new algorithms for centroiding calculation of the photon shower







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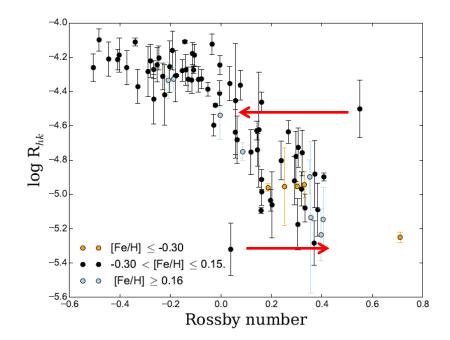


Modelado de Placas Micro-Canal como Detectores Astronómicos de Radiación UV

> Trabajo de tesis presentado por Angel Manrique Pozos Flores

Science with the WSO-UV

- Properties of exo-planetary atmospheres and the impact of UV stellar radiation in their chemistry
- Space UV features as clocks in solar-like stars
- Sinergies with large ground based infrastructures (The Large Millimeter Telescope)



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Proceedings of the International Astronautical Congress, IAC
2016
67th International Astronautical Congress, IAC 2016; Guadalajara; Mexico; 26 September 2016 through 30 September 2016; Code 126413
UV astronomy from space: On the ages of exo-worlds (Conference Paper)
Dagostino, M.C.<sup>a</sup> 🖾, Bartone, E.<sup>a</sup> 🖾, Aguilar, J.M.O.<sup>a</sup> 🖾, Aguilar, N.D.O.<sup>a</sup> 🖾, Montez, C.<sup>a</sup> 🖾, De Castro, A.I.G.<sup>b</sup> 🖾, Sachkov, M.<sup>c</sup> 🖾, Serna, B.O.<sup>d</sup> 🖾, Castillo, E.<sup>a</sup> 🖾,
Flores, A.P.<sup>e</sup> ⋈, Trujillo, L.<sup>e</sup> ⋈
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<sup>c</sup>Institute of Astronomy, Russian Academy of Sciences, Russian Federation
2016 13th International Conference on Electrical Engineering, Computing Science and Automatic Control, CCE 2016
21 November 2016, Article number 7751230
13th International Conference on Electrical Engineering, Computing Science and Automatic Control, CCE 2016; Mexico City; Mexico; 26 September 2016 through 30 September 2016
Category numberCFP16827-USB; Code 125070
Modeling micro-channel plates as astronomical detectors of UV radiation (Conference Paper)
Pozos, A.<sup>a</sup> ⋈, Castillo, E.<sup>b</sup> ⋈, Chavez, M.<sup>b</sup> ⋈, Trujillo, L.<sup>a</sup> ⋈
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Caption of Fig. Rh+k vs Rossby number

Correlation between the Rossby number (=rotational period/convection turnover time) and the photospherically corrected flux in emission of the MgII (h+k) doublet at 2800 angstrom. Fluxes were collected from IUE and HST/STIS at high resolution in a sample of ~100 F, G and K type stars on the main sequence. Such a correlation will be used to establish rotation periods and ages of other stars. Outliers correspond to objects for which we believe the measured rotational periods (through the modulation of H+K CaII lines) are wrong and the right most point corresponds to a very metal poor object.

WSO will be extremely valuable to:

- Establish a real "mean" emission MgII flux through variability monitoring
- Investigate the effect of metallicity on the correlation
- Increase the stellar sample in the UV including targets in open clusters where age determinations can be obtained through isochrone fitting.